Case studies on the effect of pre-event background in solar particle event timing

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Aug 6 2010 Turku, ECRS Parametric study Fixing the problem?

Velocity Dispersion Analysis (VDA)



A simulated SEP event (1-100 MeV protons) overlaid on a pre-event background VDA fit on a simulated event, compared to "nominal" fit

- Measure the arrival time of the first particles and fit $t_{arrival} = t_{injection} + s/v$
- Leave out "bad channels" (with max $\{I_{event}\} < 10 \cdot I_{bg}$)

Parametric study Fixing the problem?

What's wrong with the velocity dispersion?

- The event onset is masked by the pre-event background, to a varying degree
- -> The obtained "solar onset time" is wrong
 - ... but how much wrong?



Simulated solar event overlaid on a pre-event background for 5 MeV proton

Study strategy

- Simulate particle events with varying injections and transport conditions
 - Use 1-100 MeV protons
 - Overlay on varying levels of energy-dependent pre-event backgrounds
 - See how well the velocity dispersion analysis works
- Simulations based on Monte Carlo method
 - parallel transport along Parker spiral, guiding centre approximation
 - SQLT isotropic scattering off parallel Alfvén waves (method from Kocharov et al 1997)
 - adiabatic focusing and deceleration taken into account

Parametric study

- $R = I_{event}/I_{bg}$ at 88 MeV, $\Delta q < 0$ for soft event
- Left panel: the fitted pathlength (nominal 1.12 AU). The thick curve 1.4 AU, 0.2 AU contours, increasing to upper right.
- Right panel: the fitted onset time. The thick curve the nominal, 10 min contours, the green and red contour fills show delayed fit result



Velocity dispersion in high-intensity events seems to be usable

• Large, tens of minutes error for "non-optimal" events

Correction to the onset times?

- Fit rise rate, $D = d \ln(I)/dt$, to analytic estimates?
- Diffusion equation?

$$I \propto \frac{1}{(4\lambda vt/3)^{3/2}} e^{-3r^2/(4\lambda vt)}$$

gives for the rise rate $Dpprox 3\,r^2/(4\lambda\,v\,t^2)$ at the onset

- First particles arrive to 1 AU at t = ε: not usable at the onset
- Lacks physics relevant for the onset
- Telegraph equation?
 - "Signal speed" v_s : First particles arrive at $t = r/v_s$
 - Tricky, Bessel functions, no solutions with adiabatic changes

Simple estimate

- Use onset error time $t' = t r/v_{particle}$ in diffusion equation
- Focusing, deceleration and scattering model modify the arrival time: Fit the modification from the simulated events

> Rise rate

$$D \equiv \frac{\partial \ln l}{\partial t'} = A \frac{3}{4} \frac{r^2}{\lambda v t'^2}$$



-> obtain correction
$$t' = \sqrt{A \frac{3}{4} \frac{r^2}{\lambda v t D}}$$
 for all energy channels

Simple estimate



• Right-hand panel: VDA with corrected onset times, using λ =0.3 AU for all energies

July 25, 1997

- Fairly steep rise, background and maximum spectra similar
- Uncorrected onset 20:11, corrected to 20:01
- Uncorrected pathlength 2.35, corrected to 2.18



CME at 19:54, 611 km/s, western, wide

Parametric study Fixing the problem?

Aug 20 2002

- Significantly masked at low energies
- Uncorrected onset 08:30, corrected to 08:26
- Uncorrected pathlength 1.07, corrected to 1.05



8:22-8:26-8:30 M3.4 S10W38 Southern partial-halo CME 8:10, 1099 km/s, at 4.9 rs at 8:54

Aug 25 2005

- Gradual rise, masked at low energies
- Uncorrected onset 12:22, unchanged
- Uncorrected pathlength 2.01, corrected to 1.7



4:30-4:40-4:45 M6.4 N09E80, CME at 4:16, 1327 km/s (eastern partial halo)

Conclusions

- Velocity dispersion analysis suffers from pre-event background: large errors possible
- Large events ok, partially masked events may have large errors
- Corrections based on rise rate can be devised
- Won't correct "everything"
 - More physics needed in analysis
 - All assumptions not satisfied (e.g. Parker field, simultaneous injection)